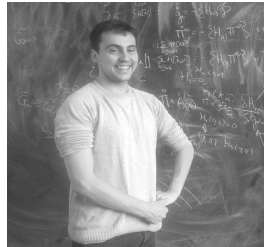
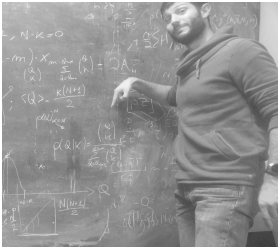
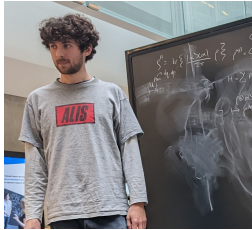


# Decoherence vs Diffusion: Testing the Quantum Nature of Spacetime



WITH

A. Grudka, T. Morris, Muhammad Sajjad, Andrea Russo, Zach Weller-Davies, Barbara Soda, Carlo Sparaciari, Emanuele Panella, A. Pontzen

Phys. Rev. X 13, 041040 (2023)

Nature Comms 14, 7910 (2023)

arXiv:2302.07283

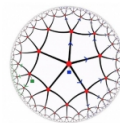
arXiv:2402.17844

**Jonathan Oppenheim**

**@postquantum**

ISQIQC Kokata

31 March 2025



**It from Qubit**

Simons Collaboration on  
Quantum Fields, Gravity and Information



**Engineering and  
Physical Sciences  
Research Council**



**UCL**

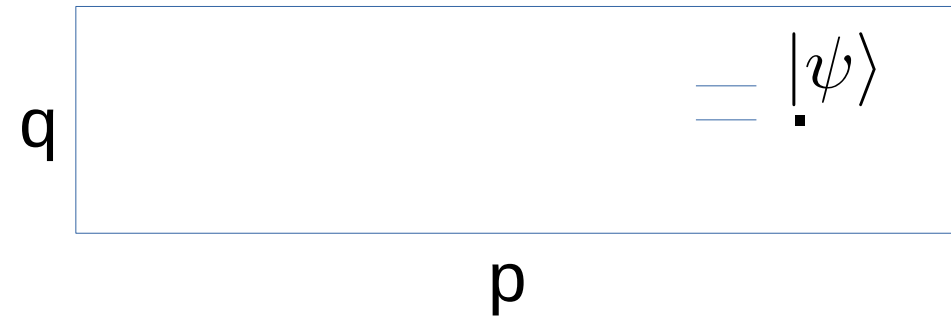
# Frameworks

## Quantum Mechanics

$$\frac{\partial \hat{\sigma}}{\partial t} = -i[\hat{H}, \hat{\sigma}]$$

## Classical Mechanics

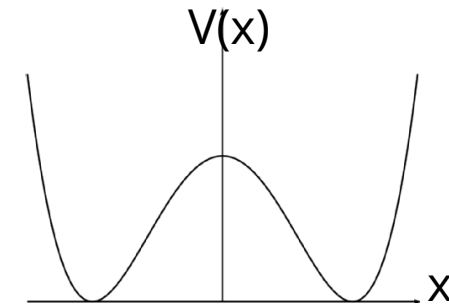
$$\frac{\partial \rho(q,p)}{\partial t} = \{H(q,p), \rho(q,p)\}$$



# Frameworks

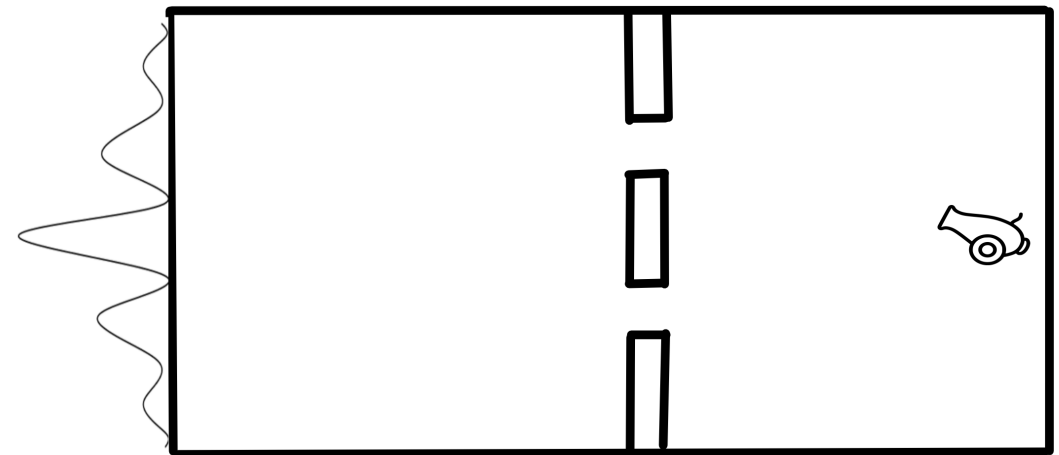
## Quantum Mechanics

$$\frac{\partial \hat{\sigma}}{\partial t} = -i[\hat{H}, \hat{\sigma}]$$



## Classical Mechanics

$$\frac{\partial \rho(q,p)}{\partial t} = \{H(q,p), \rho(q,p)\}$$



**Back-reaction**



# Motivation: gravity is not a gauge theory

---

Gauge transformations  
act on  $\Psi(x)$ , Diffs move  $x$

---

Universal geometry on  
which fields live

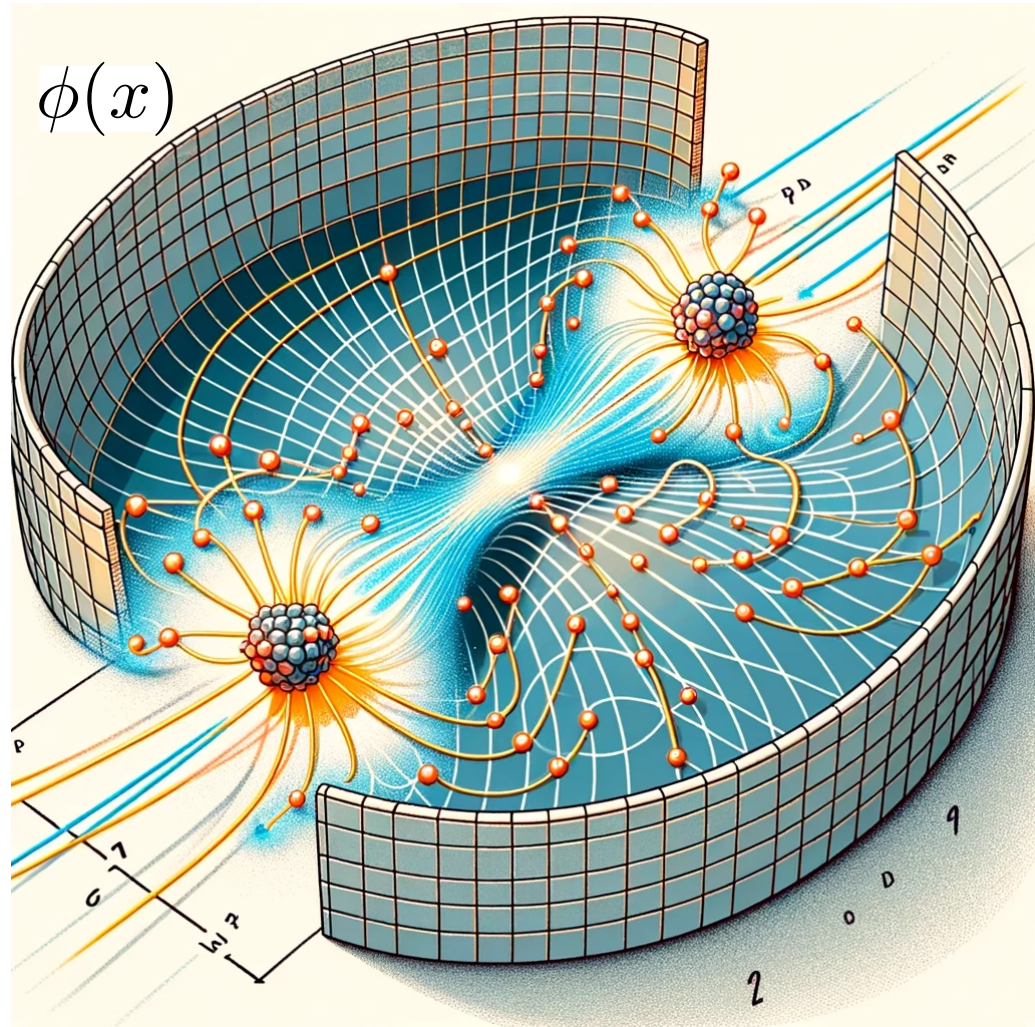
---

Wheeler-deWitt Eqn

$$\mathcal{H}(x)|\psi\rangle = 0$$

---

Non-renormalisable  
(quantum theory)



# Classical-quantum gravity

If true

---

Formally  
renormalisable

---

Black hole information  
problem?

---

Small  $\Lambda$ ?  
dark matter?

EXPERIMENT

---

Born rule & No need for  
Measurement postulate

---

Decoherence vs  
Diffusion trade-off

EXPERIMENT

# Classical-quantum dynamics

## Debate

---

### No

Feynman (1957)  
DeWitt (1962)  
Unruh (1984)  
Aharonov (~1986)  
Eppley & Hannah (1977)  
Unruh (1984)  
Caro & Salcedo (1999)  
Terno (2004)  
Carlip (2008)  
Marletto Vedral (2017)  
Galley, Giacomini, Selby (2022)

---

### Maybe

Sherry & Sudarshan (1978)  
Boucher & Traschen (1988)  
Kapral (1999)  
Peres & Terno (2001)  
Hall & Reginatto (2005)  
Mattingly (2006)  
Albers, Kiefer & Reginatto (2008)  
Kent (2018)

# Classical-quantum dynamics

## History

---

Semi-classical Einstein  
(pathological when  
fluctuations are large)

Page & Geilker (1981);  
Gisin (1989)

---

Simple examples

Blanchard & Jadczyk (1994);  
Diosi (1995);  
Poulin (2017);

Kafri, Taylor, Milburn (2014);  
Diosi, Tilloy (2016)

---

Quantum chemistry  
(negative probabilities)

Kapral review (2006);  
Koopman-von Neumann (1931-32)

---

Experiments!

Bose et. al. (2017);  
Marletto et. al. (2017)  
Lami, Pedarnals, Plenio (2022)  
Carney (2108.06320)

# A post-quantum theory of classical gravity?

---

What is the most general form of CQ dynamics?

---

---

Decoherence vs Diffusion: testing quantum gravity

---

---

Renormalisable without Ghosts!

---

---

Anomalous contribution to the metric (dark matter, dark energy?)

---

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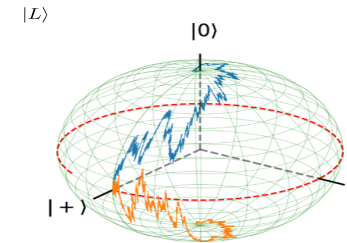
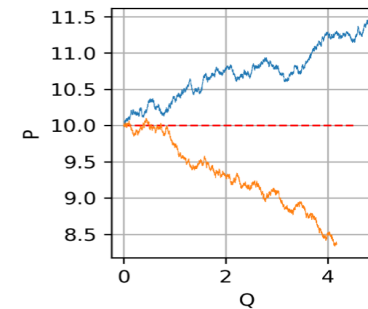
# CQ Dynamics

## Path Integral

$$\rho(q, p, \phi^\pm, t_f) = \int \mathcal{D}q \mathcal{D}p \mathcal{D}\phi^\pm e^{iS_C[q, p] + iS[\phi^+] - iS[\phi^-] + iS_{FV}[\phi^\pm] + iS_{CQ}[q, p, \phi^\pm]} \delta\left(\dot{q} - \frac{p}{m}\right) \rho(q, p, \phi^\pm, t_i)$$

JO, Zach Weller-Davies (2023)

## Trajectories

JO, I. Layton, Z. Weller-Davies (2022)  
Tilloy (2024)

## Master Eqn

$$\begin{aligned} \frac{\partial \hat{\rho}}{\partial t} \approx & \{H^{(grav)}, \hat{\rho}\} - i[\hat{H}^{(m)}, \hat{\rho}] + \frac{1}{2}\{\hat{H}^{(m)}, \hat{\rho}\} - \frac{1}{2}\{\hat{\rho}, \hat{H}^{(m)}\} \\ & + \int dx dx' \frac{\delta^2}{\delta\pi_\Phi(x)\delta\pi_\Phi(x')} (D_2(x, x')\hat{\rho}) + \frac{1}{2} \int dx dx' D_0(x, x') ([\hat{m}(x), [\hat{\rho}, \hat{m}(x')]]) \end{aligned}$$

CPTP MAP

JO, Sparaciari, Soda, Weller-Davies

# Frameworks

## Quantum Mechanics

$$\frac{\partial \hat{\sigma}}{\partial t} = -i[\hat{H}, \hat{\sigma}]$$

$$q \quad \boxed{\quad \quad \quad = |\psi\rangle} \quad p$$

## Classical Mechanics

$$\frac{\partial \rho(q,p)}{\partial t} = \{H(q,p), \rho(q,p)\}$$

$$\hat{\rho}(q,p) = \rho(q,p) \begin{pmatrix} p(0|q,p) & \alpha(q,p) \\ \alpha^*(q,p) & p(1|q,p) \end{pmatrix}$$

# Classical, quantum, & CQ States

Q

HILBERT SPACE

 $\hat{\sigma}$ 

$$\text{tr } \hat{\sigma} = 1$$

POSITIVE MATRIX

C

PHASE SPACE

 $\rho(q, p)$ 

$$\int dq dp \rho(q, p) = 1$$

POSITIVE DISTRIBUTION

CQ

$$\hat{\rho}(q, p) = \rho(q, p) \hat{\sigma}(q, p)$$

$$\int dq dp \text{tr } \hat{\rho}(q, p) = 1$$

POSITIVE MATRIX AT EACH POINT

# Classical, quantum, & CQ States

Q

HILBERT SPACE

$$\text{tr } \hat{\sigma} = 1 \quad \hat{\sigma} = \begin{pmatrix} p(0) & \alpha \\ \alpha^* & p(1) \end{pmatrix}$$

POSITIVE MATRIX

C

PHASE SPACE

$$\int dq dp \rho(q, p) = 1$$

POSITIVE DISTRIBUTION

CQ

$$\hat{\sigma}_{cq} = \int dz \rho(z; t) |z\rangle\langle z| \otimes \sigma(z; t)$$

$$\hat{\rho}(z) = \rho(z) \begin{pmatrix} p(0|z) & \alpha(z) \\ \alpha^*(z) & p(1|z) \end{pmatrix}$$

$$\int dz \text{tr } \hat{\rho}(z) = 1$$

POSITIVE MATRIX AT EACH Z

# Dynamics must be linear and preserve state-space

$$\mathbb{1} \otimes \mathcal{L}$$

Must be positive

Norm preserving

positive matrix at each q,p

$$\int dq dp \text{tr} \hat{\rho}(q, p) = 1$$

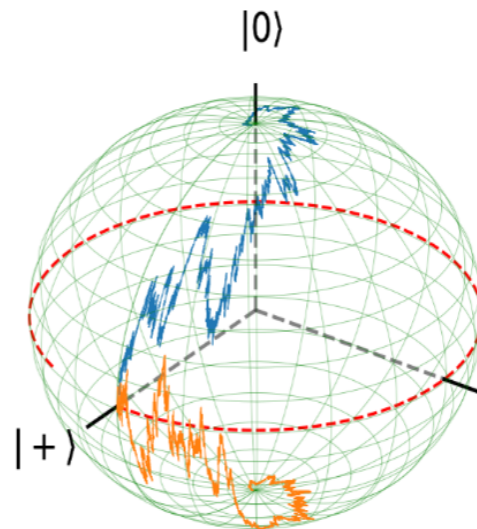
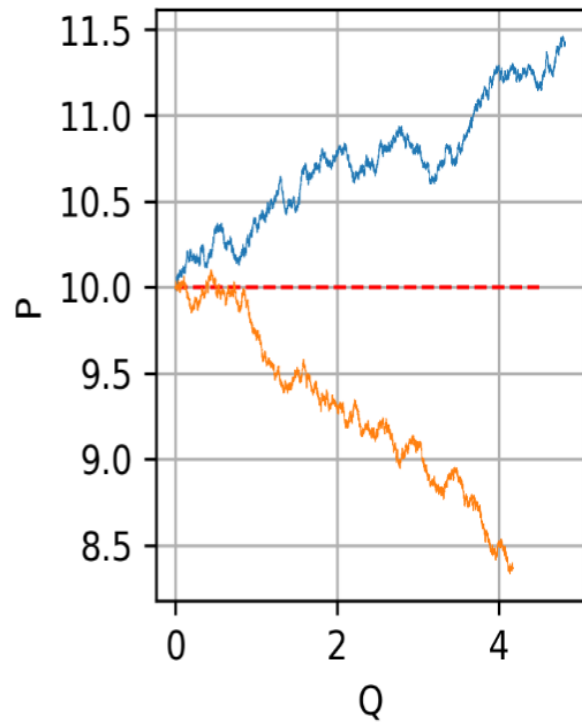
$$|\psi\rangle = \frac{1}{\sqrt{2}} (|0\rangle|\psi_L\rangle - |1\rangle|\psi_R\rangle) \quad \sigma^{(m)} = \frac{1}{2} |\psi_L\rangle\langle\psi_L| + \frac{1}{2} |\psi_R\rangle\langle\psi_R|$$

$$\frac{1}{2} \mathcal{L}(|\psi_L\rangle\langle\psi_L|) + \frac{1}{2} \mathcal{L}(|\psi_R\rangle\langle\psi_R|) = \mathcal{L}\left(\frac{1}{2} |\psi_L\rangle\langle\psi_L| + \frac{1}{2} |\psi_R\rangle\langle\psi_R|\right)$$

# Example of continuous master-equation

**Stern-Gerlach**

$$\hat{H} = \frac{p^2}{2m} + D_1 q \hat{\sigma}$$

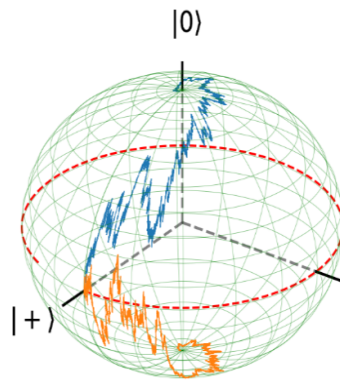
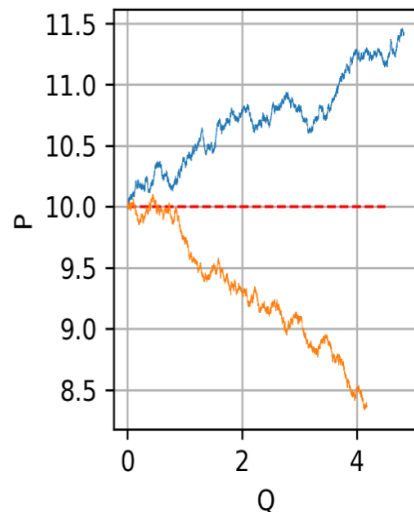


$$D_2 D_0 \succeq D_1^2$$

# Example of continuous master-equation

**Stern-Gerlach**  $\hat{H} = \frac{p^2}{2m} + D_1 q \hat{\sigma}$

$$\frac{\partial \hat{\rho}}{\partial t} = -i[\hat{H}, \hat{\rho}] + \frac{1}{2}\{\hat{H}, \hat{\rho}\} - \frac{1}{2}\{\hat{\rho}, \hat{H}\} + \frac{D_2}{2}\{q, \{q, \hat{\rho}\}\} + \frac{D_0}{2}[\hat{\sigma}, [\hat{\rho}, \hat{\sigma}]]$$



$$D_2 D_0 \succeq D_1^2$$

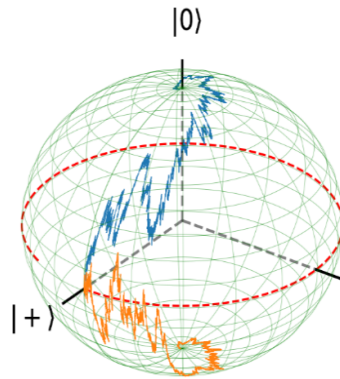
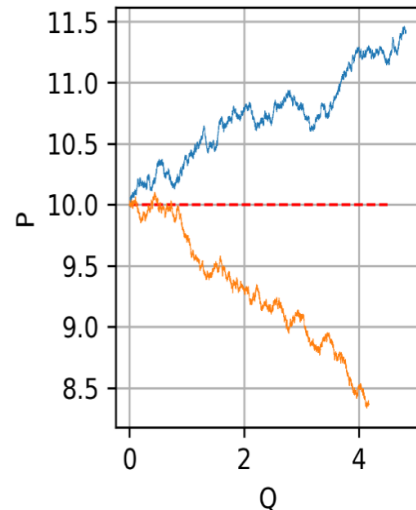
Diosi (1995)

Isaac Layton, JO, Zach Weller-Davies (2022)

# Example of continuous master-equation

**Stern-Gerlach**  $\hat{H} = \frac{p^2}{2m} + D_1 q \hat{\sigma}$

$$\frac{\partial \hat{\rho}}{\partial t} = -i[\hat{H}, \hat{\rho}] + \frac{D_1}{2} \left( \hat{\sigma} \frac{\partial \hat{\rho}}{\partial p} + \frac{\partial \hat{\rho}}{\partial p} \hat{\sigma} \right) + \frac{D_2}{2} \frac{\partial^2 \hat{\rho}}{\partial p^2} + \frac{D_0}{2} [\hat{\sigma}, [\hat{\rho}, \hat{\sigma}]]$$



$$D_2 D_0 = D_1^2$$

Diosi (1995)

Isaac Layton, JO, Zach Weller-Davies

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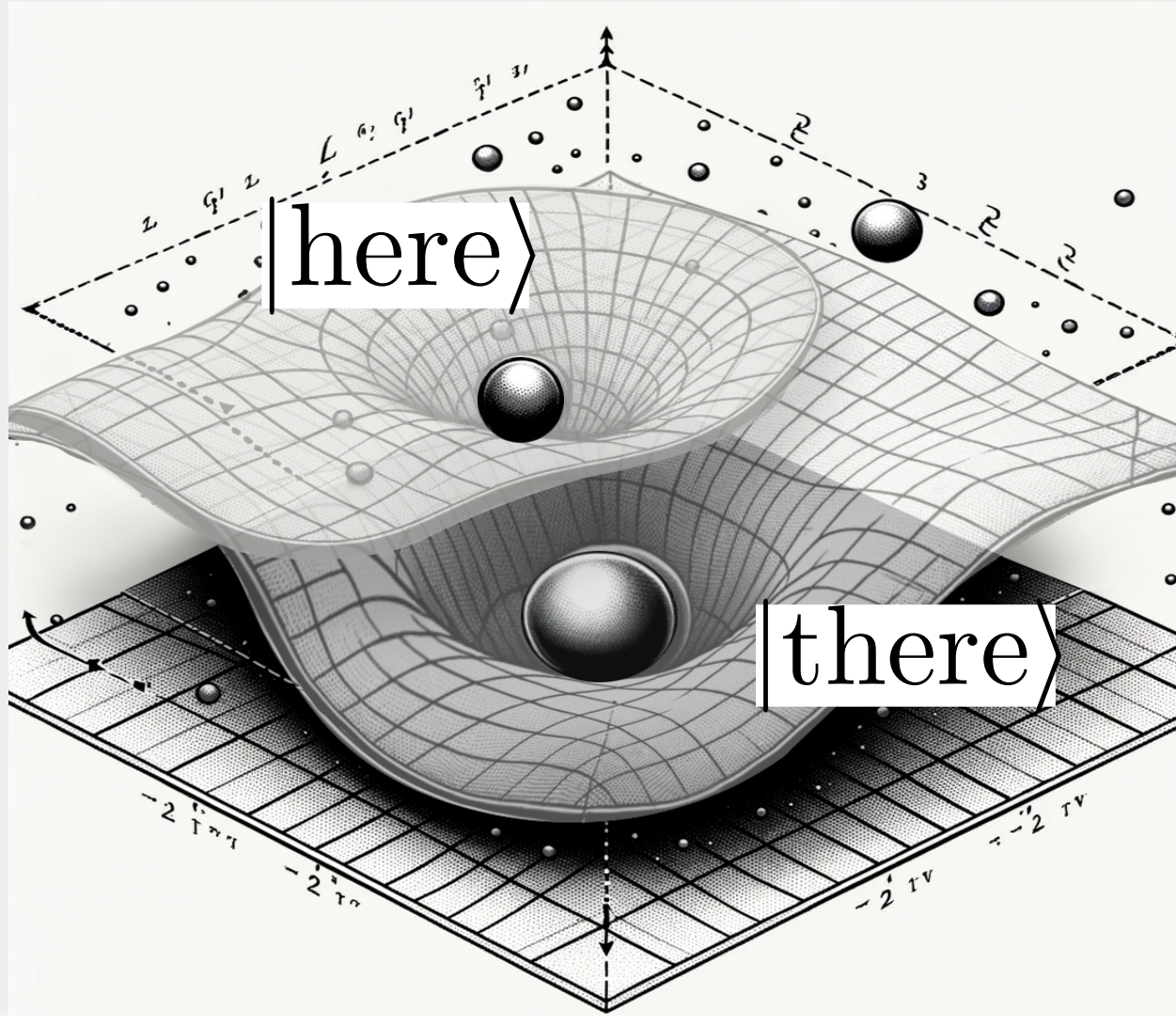
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Anomalous contribution to the metric (dark matter, dark energy?)

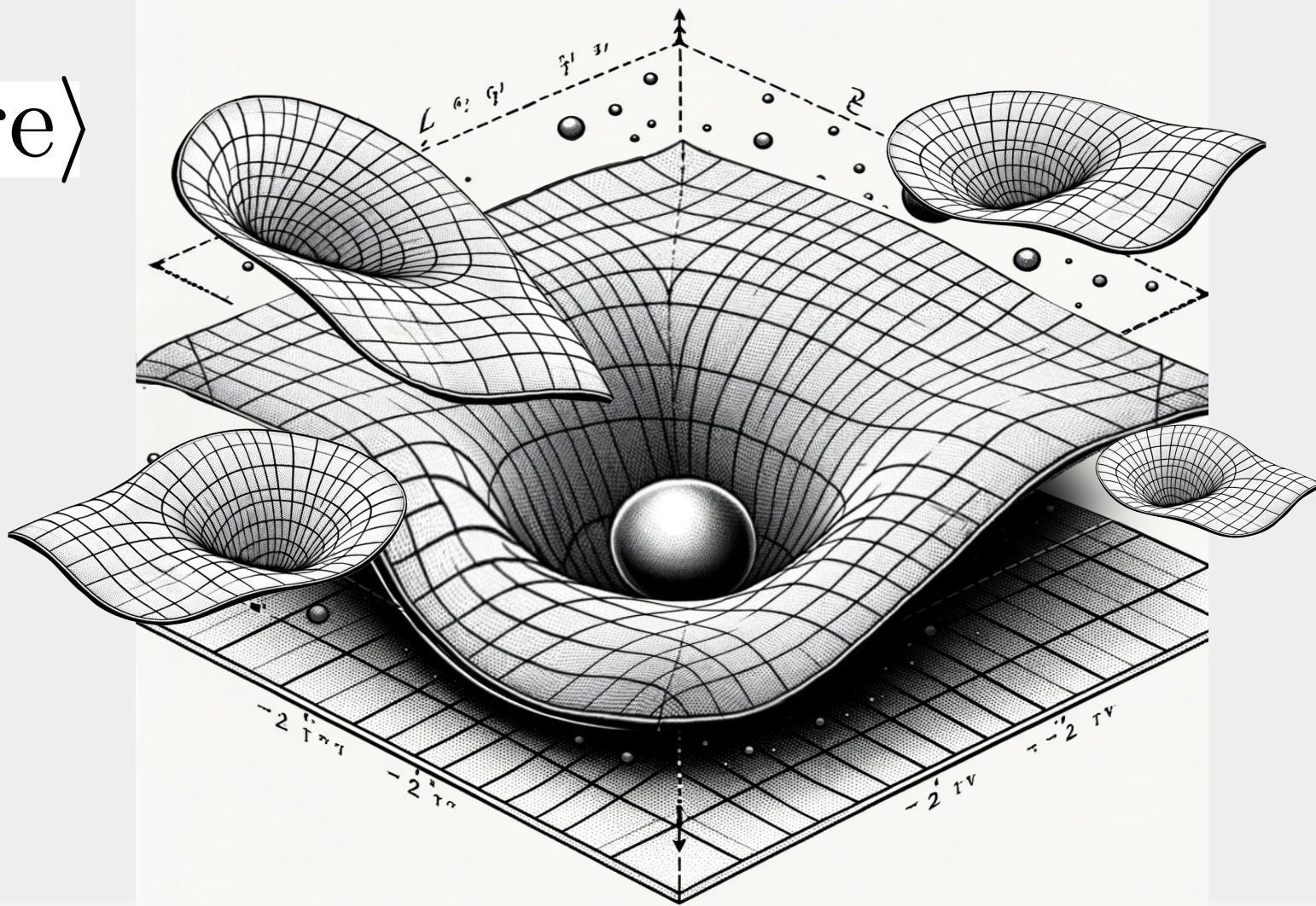
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# Decoherence vs diffusion trade-off



# Decoherence vs diffusion trade-off

$| \text{here} \rangle$

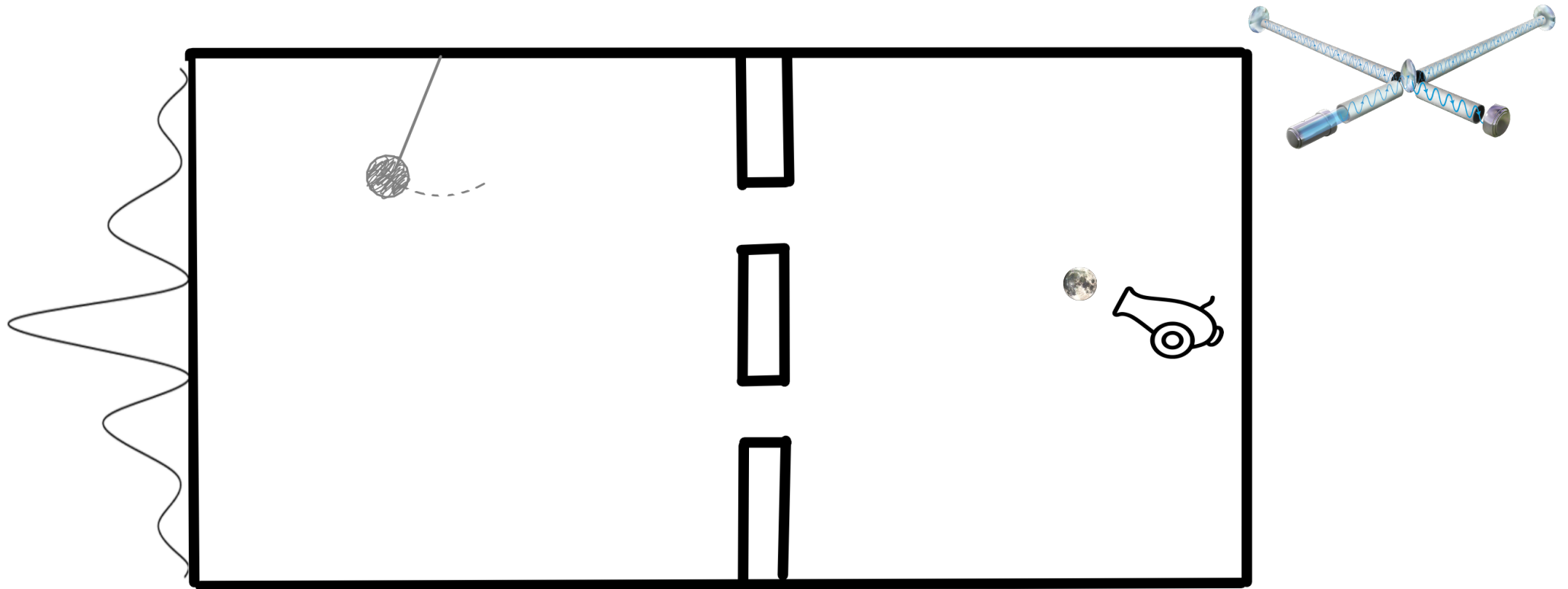


$| \text{there} \rangle$



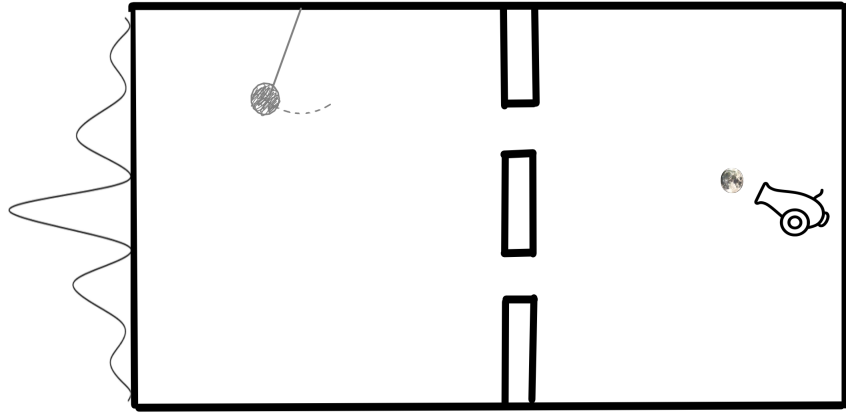


# Decoherence vs diffusion



$$|\psi\rangle = \frac{1}{\sqrt{2}} (|E_L\rangle |L\rangle + |E_R\rangle |R\rangle)$$

# Decoherence vs diffusion

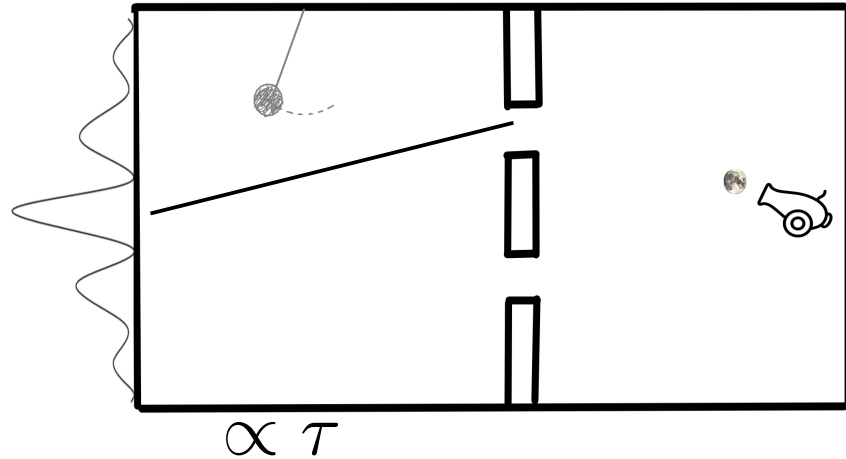


$$\hat{\sigma}(t) = \begin{pmatrix} \frac{1}{2} & \alpha^*(t) \\ \alpha(t) & \frac{1}{2} \end{pmatrix}$$

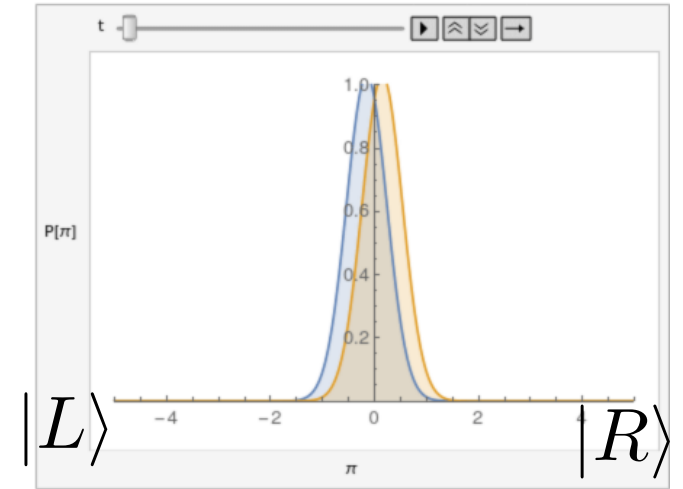
$$\alpha(t) = \langle E_L(t) | E_R(t) \rangle$$

$$|\psi\rangle = \frac{1}{\sqrt{2}} (|E_L\rangle |L\rangle + |E_R\rangle |R\rangle)$$

# Decoherence vs diffusion



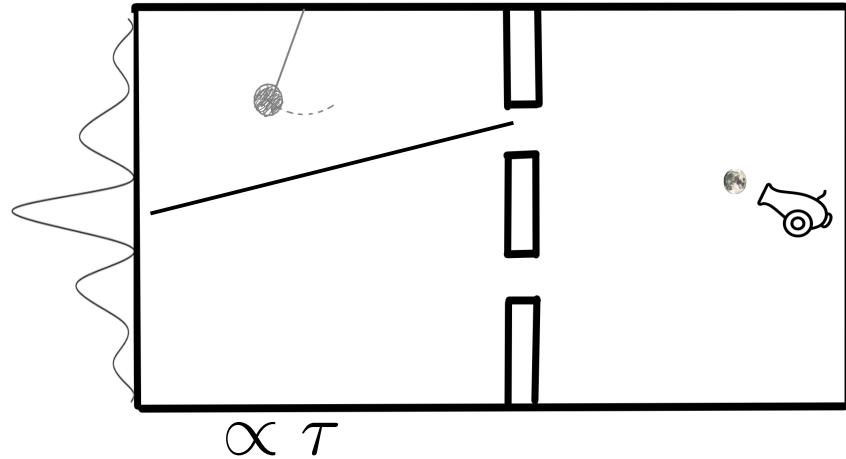
$$\tau \leq \frac{D_2}{\langle F \rangle^2}$$



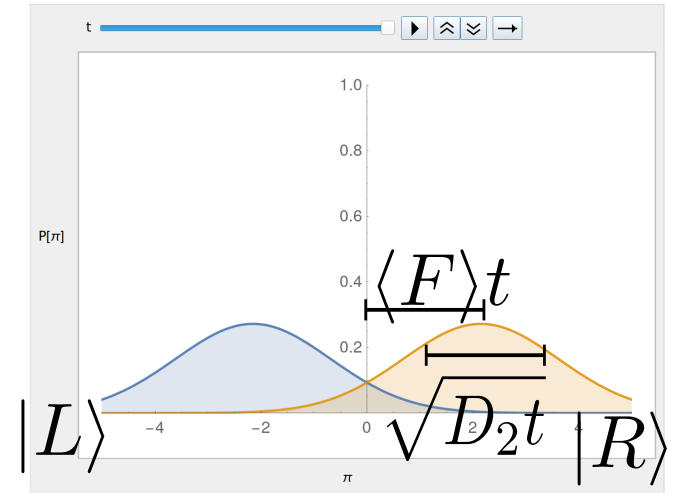
$$D_2(q, p) - D_1^\dagger(q, p) D_0^{-1}(q, p) D_1(q, p) \succeq 0$$

Holds for all classical-quantum dynamics

# Decoherence vs diffusion



$$\tau \leq \frac{D_2}{\langle F \rangle^2}$$



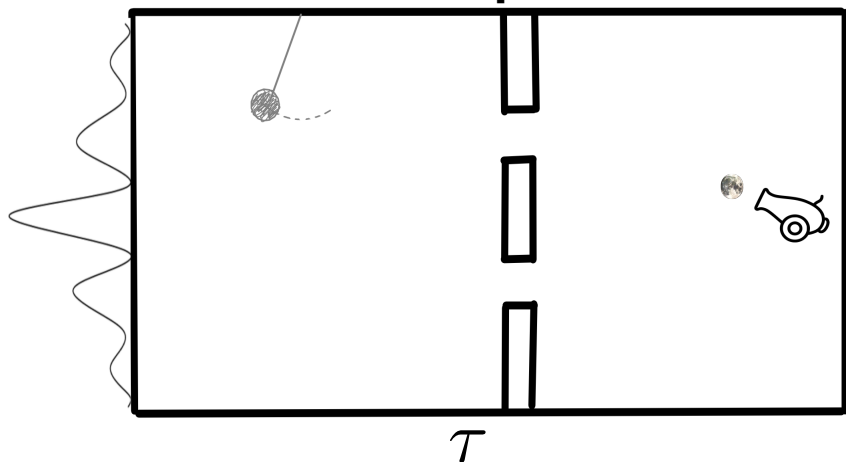
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Holds for all classical-quantum dynamics

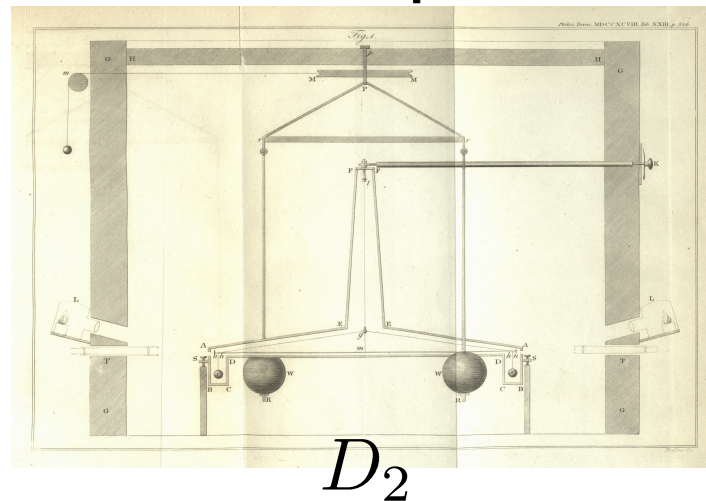


# Decoherence vs diffusion

## Double Slit Experiment



## Cavendish Experiment



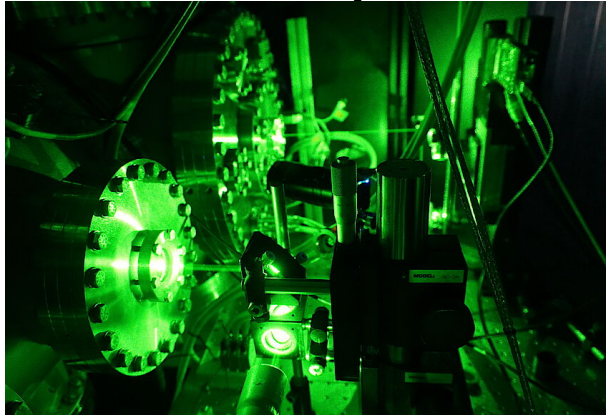
$$\langle F(x) \rangle = \langle \hat{m}(x) \rangle$$

$$\tau \leq \frac{D_2}{\langle F \rangle^2}$$



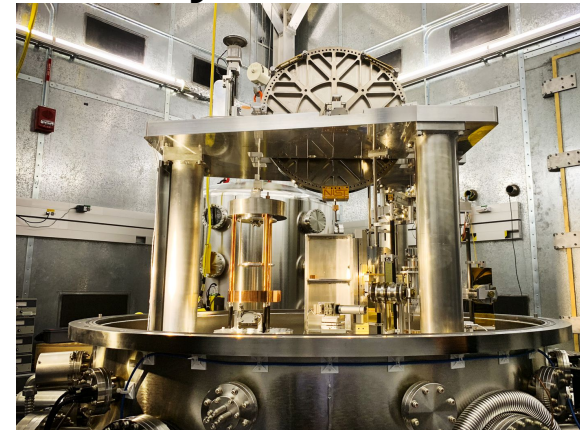
# Figures of merit

**Double Slit Experiment**



$\tau$

**Gravity measurement**



$D_2$

$$\tau \leq \frac{D_2}{\langle F \rangle^2}$$

$$\tau \approx \frac{D_G c^3}{G_N^2} \frac{V}{M^2} \begin{array}{|l|l|} \hline \text{Master Equation} & \text{Diffusive} \\ \hline \text{Continuous (ultra-local)} & D_2(\Phi; x, y) = D_2(\Phi)\delta(x, y) \\ & D_2(\Phi) = \sum_n c^n \Phi^n \quad \left| 10^{-41} \geq D_2 \geq 10^{-9} \text{ kg}^2 \text{sm}^{-3} \text{ (Eqn (44))} \right. \\ \hline \text{Continuous (Eqn (D11) or (D13))} & D_2(\Phi; \vec{x}, \vec{x}'; \omega) \Big|_{\omega_0 \rightarrow 0} = \frac{D_G c^3}{4\pi} \left( \frac{\omega^3}{\omega_0^2 c} + \frac{1}{|x - x'|} \right) \\ \hline \text{Discrete (ultra-local)} & D_2(\Phi; \vec{x}, \vec{x}'; \omega) \Big|_{\omega_0 \rightarrow 0} = \frac{D_G c^3}{4\pi} \left( \frac{\omega^3}{\omega_0^2 c} + \frac{1}{|x - x'|} \right) \\ \hline \end{array} S_{aa}(\vec{x}, \vec{x}'; \omega) := \langle a(x)a(x') \rangle - \langle a(x) \rangle \langle a(x') \rangle$$



